

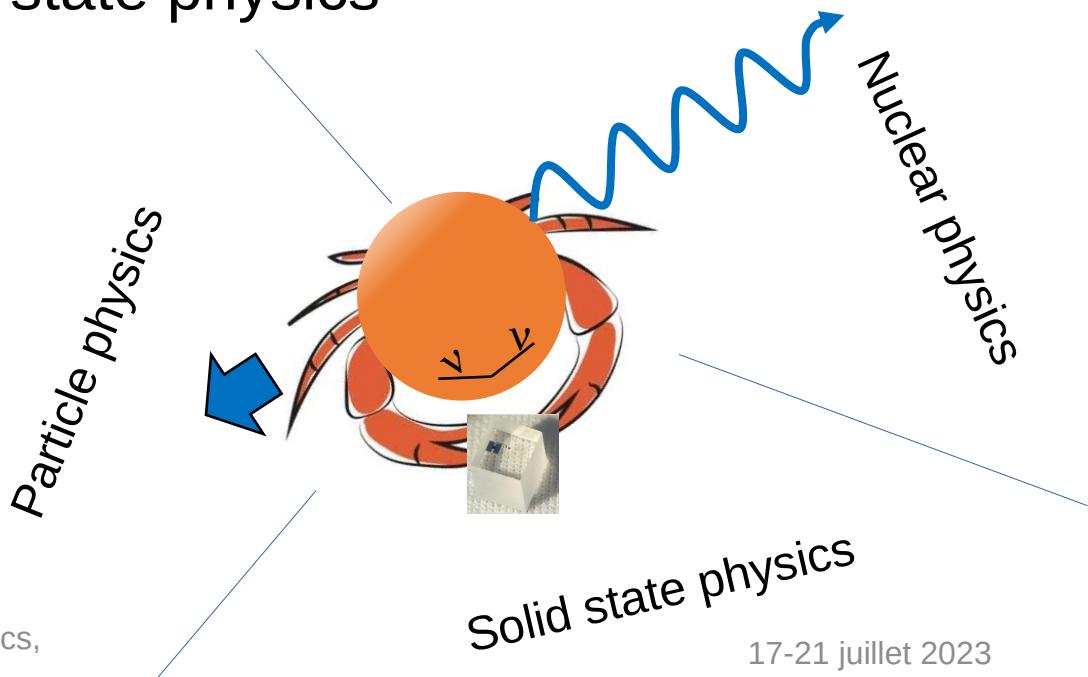


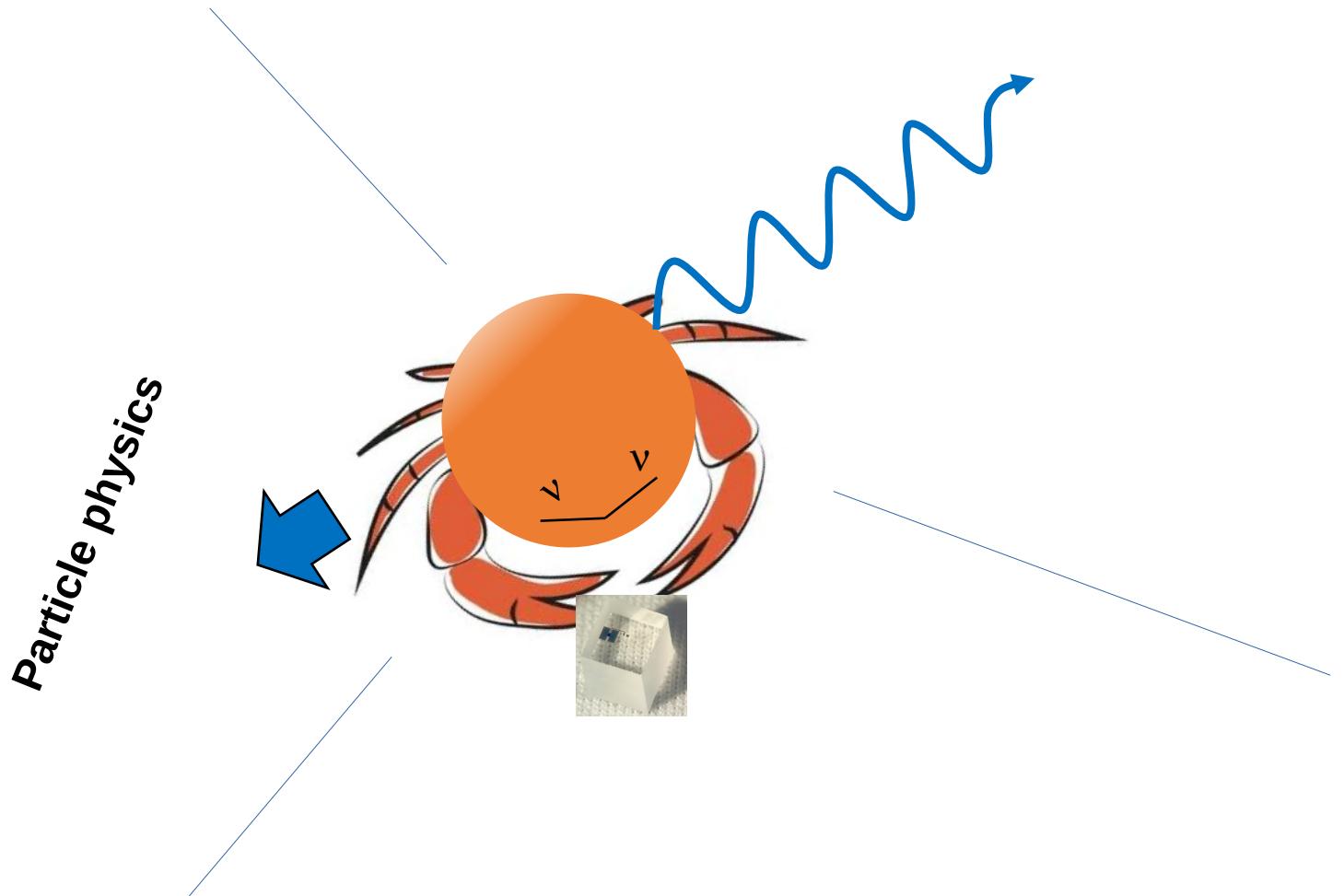
CRAB (Calibrated Recoils for Accurate Bolometry at the 100 eV scale) on the shore of particle, nuclear and solid state physics

L. Thulliez on behalf of the CRAB collaboration

CEA-Saclay/DRF/Irfu/DPhN

loic.thulliez@cea.fr

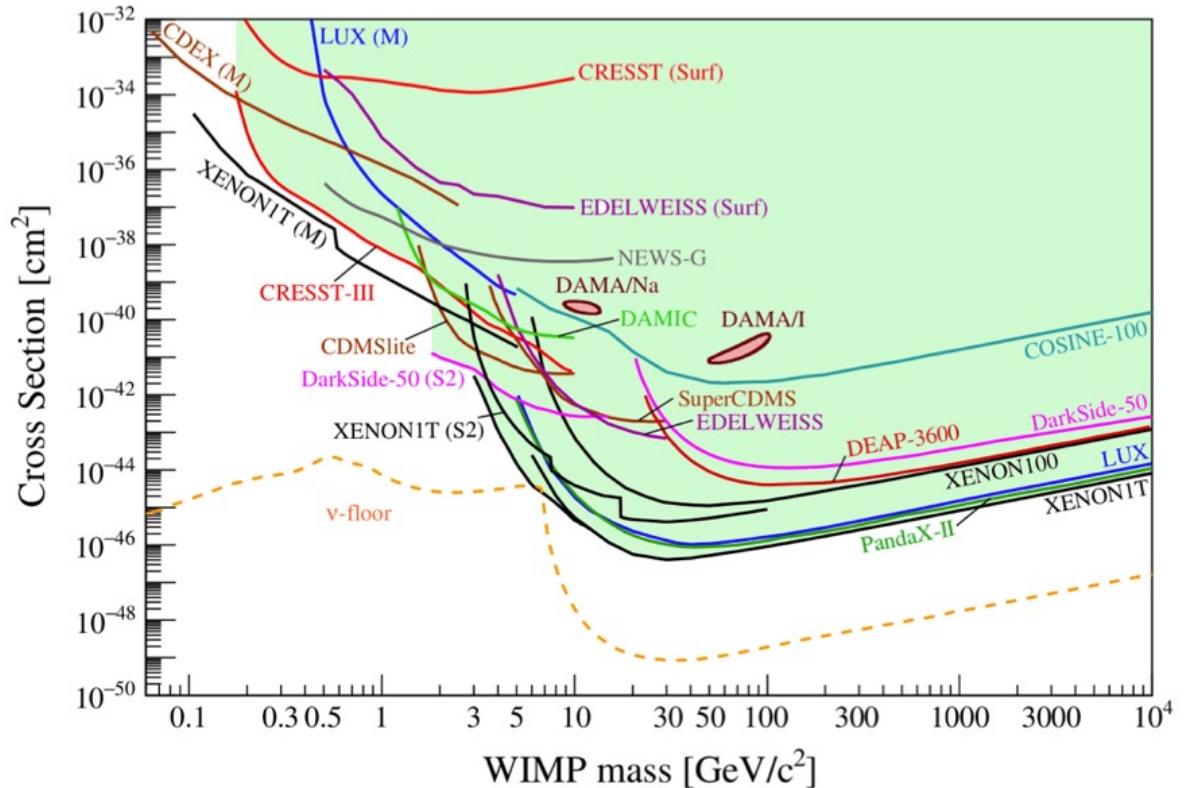




Particle physics



NEW PHYSICS WITH DARK MATTER ?



Direct Detection of Dark Matter APPEC Committee Report (2021)

- Moving to lower mass range
⇒ detection of sub-keV nuclear recoil energy
- Sensitivity in large mass range approaching the neutrino-floor limit

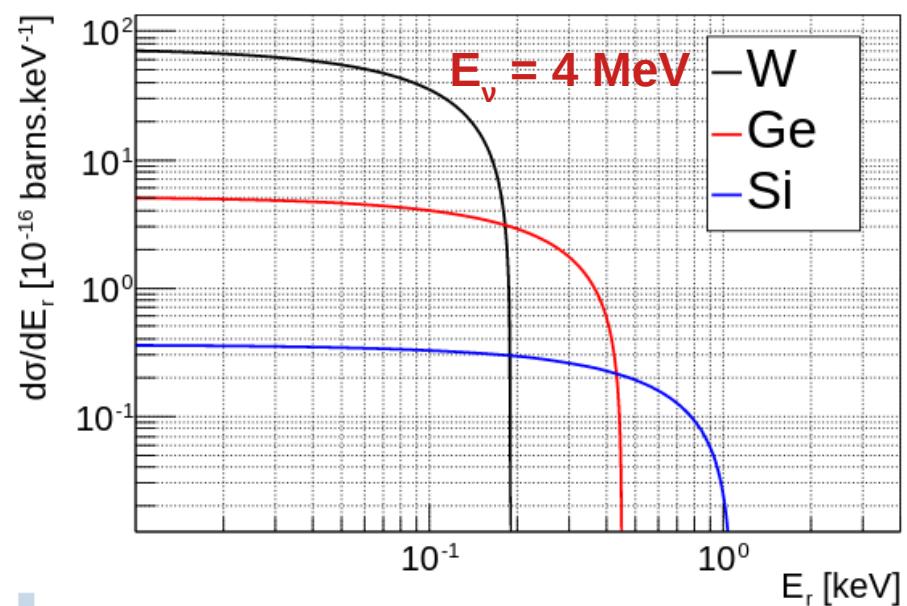
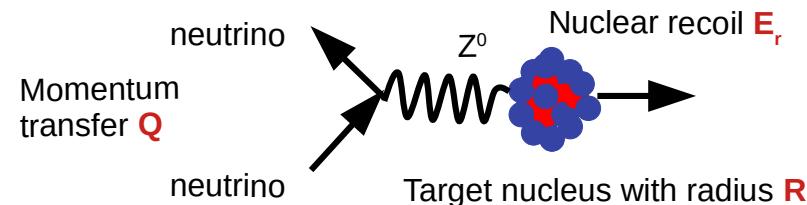


COHERENT ELASTIC NEUTRINO-NUCLEUS SCATTERING (CEvNS)

- Proposed by D. Z. Freedman in 1974 [1]
- First detected by COHERENT in 2017 [2]
- Not sensitive to neutrino flavors



- test SM (θ_w , etc)
- test BSM physics (neutrino magnetic moment, non-standard interaction, etc)
- Measure nuclear form factor (neutron skin)
- Nuclear reactor monitoring



$$\frac{d\sigma}{dE_r} = \frac{G_F^2}{4\pi} [N - Z(1 - 4\sin^2\theta_w)]^2 F^2(Q^2) M \left(1 - \frac{ME_r}{2E_\nu}\right)$$

Weak nuclear charge
 $\sin^2(\theta_w) \sim 0,231$
 $\rightarrow N^2$

Nuclear form factor
 Full coherency $QR < 1$
 Fulfill for $E_\nu < 30 \text{ MeV}$
 $\rightarrow 1$

Kinematics

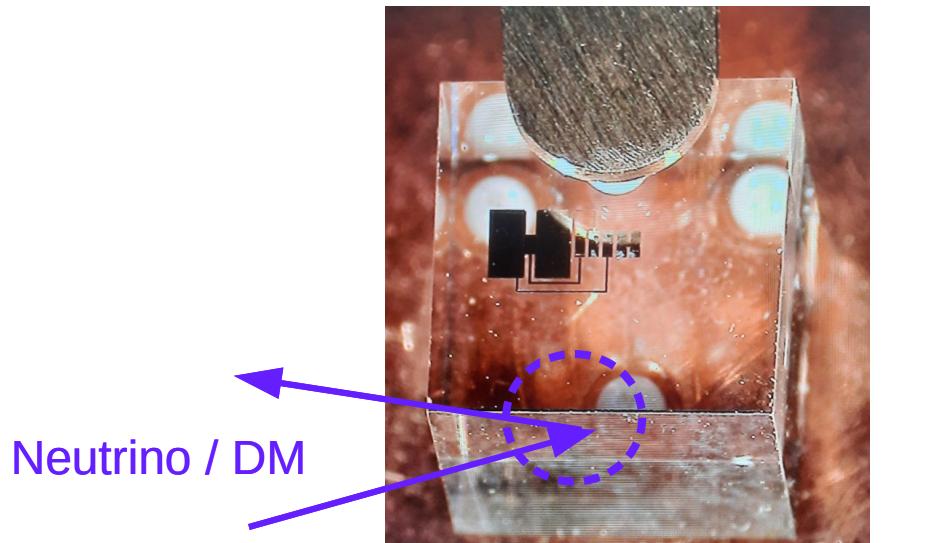
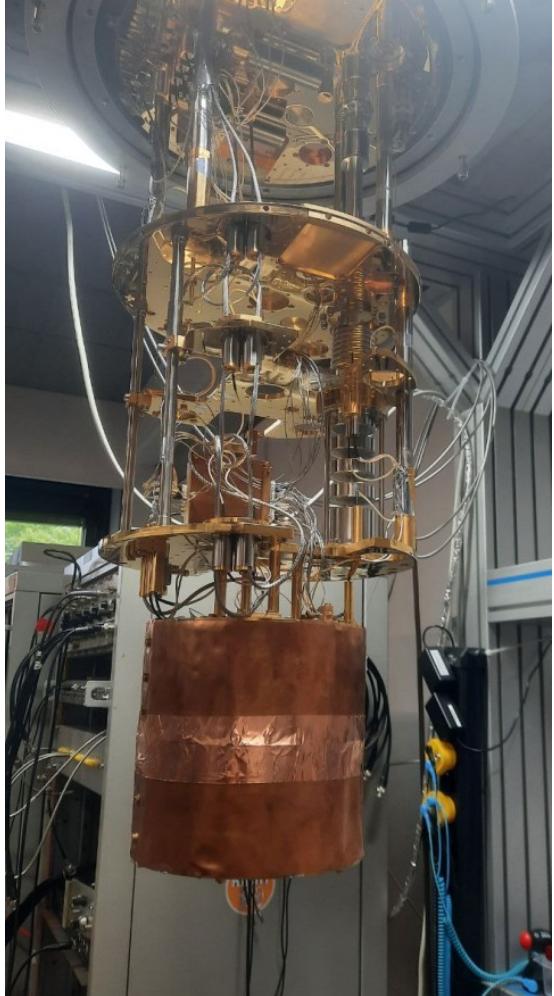
Reactor anti-neutrinos induce sub-keV energy
 \Rightarrow real challenge from the detector side
 \rightarrow good energy resolution \sim few eV
 \rightarrow low energy threshold \sim few 10 eV

Examples :
 RICOCHET @ILL [3]
 NUCLEUS @Chooz nuclear power plant [4]

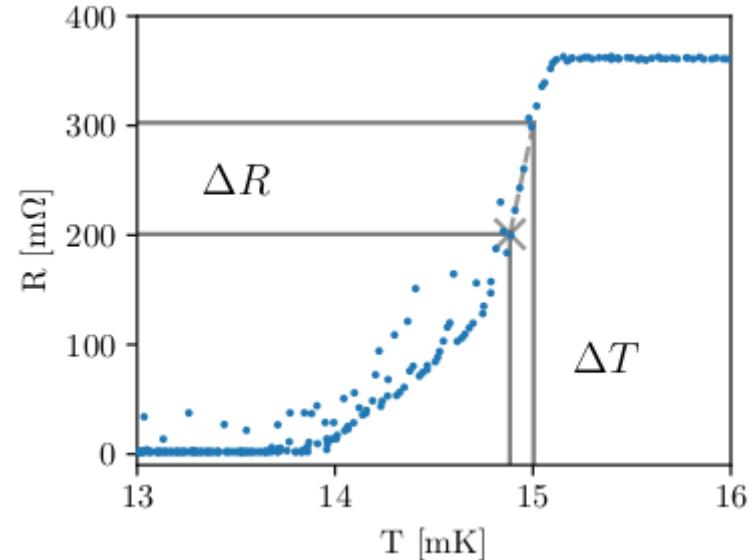


HOW TO DETECT LOW ENERGY NUCLEAR RECOILS ?

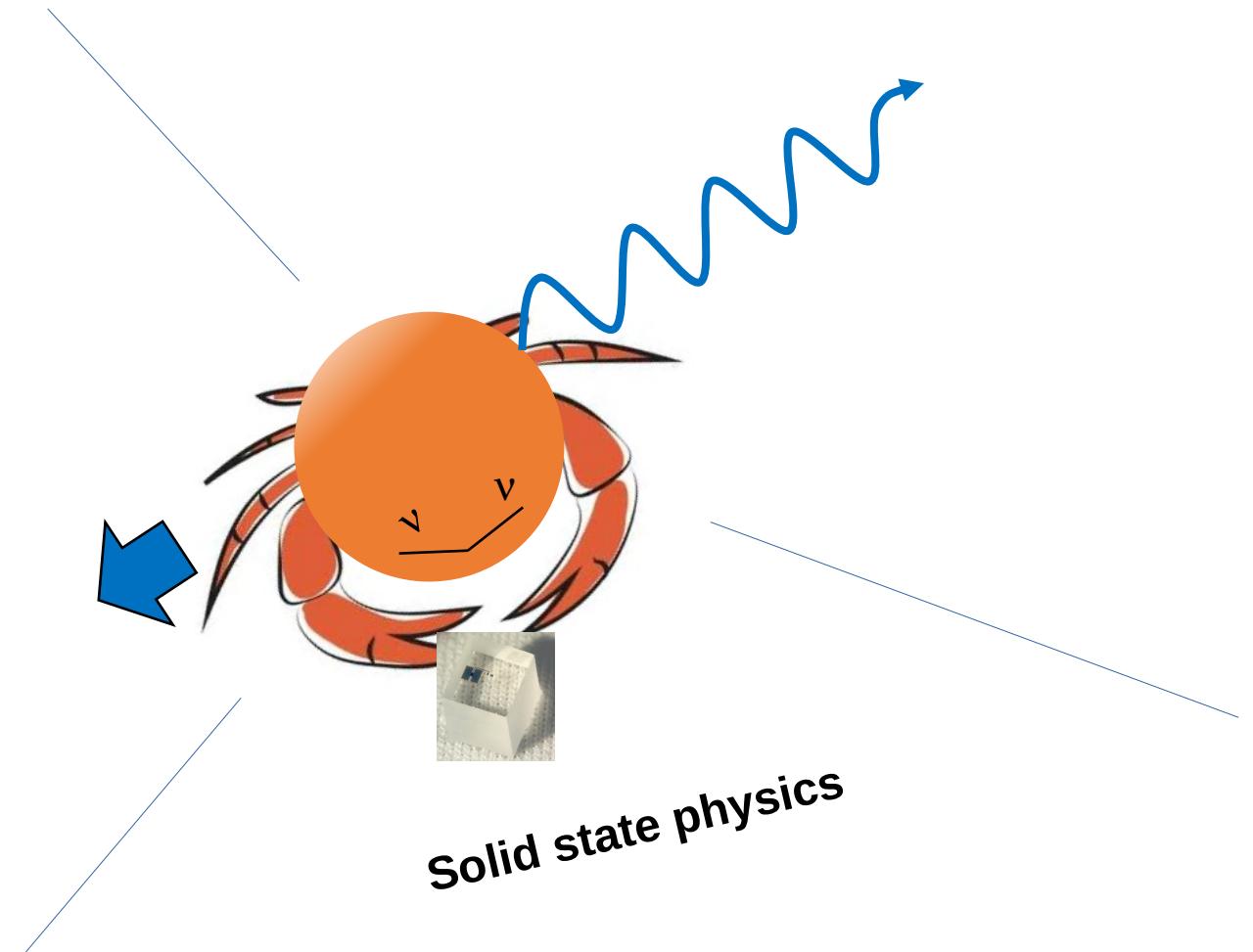
Use of **cryogenic detector** operated at **few mK**



CaWO₄ 5x5x5 mm³ NUCLEUS cryogenic detector
 $E_{th}=50$ eV $\sigma(E)=6$ eV [1]



TES resistance as a function of temperature [2]

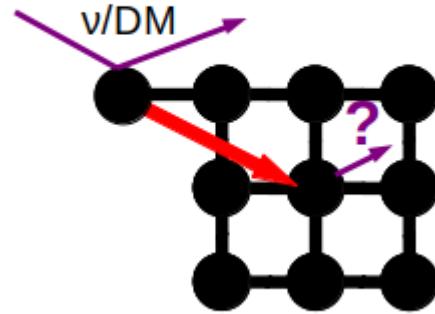


Solid state physics

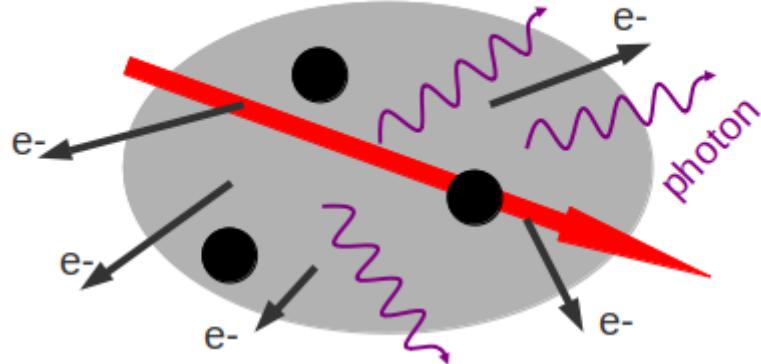
WHAT HAPPEN AFTER A PRIMARY RECOIL IN THE DETECTOR ?

Complex solid state physics to understand for precise measurements

Initial interaction



Ionization / Light emission

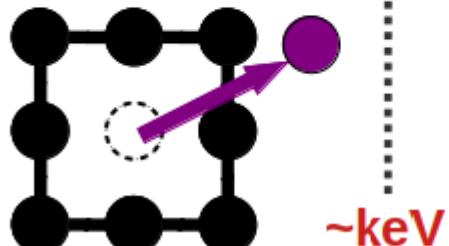


→ Quenching factor k below 1 keV ?
 $k = E_{nr}^{ioni}/E_{e-}^{ioni}$

Atomic displacement

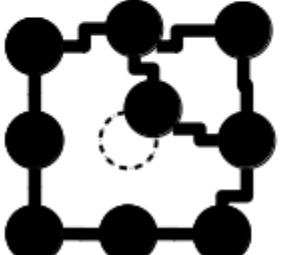
Secondary recoil

cascade



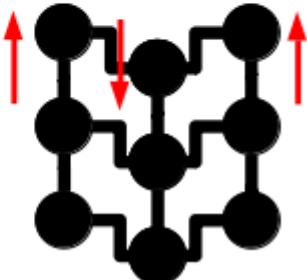
Lattice defect creation

energy trapped in the detector
 ⇒ energy biasing



Phonon excitation

temperature increase



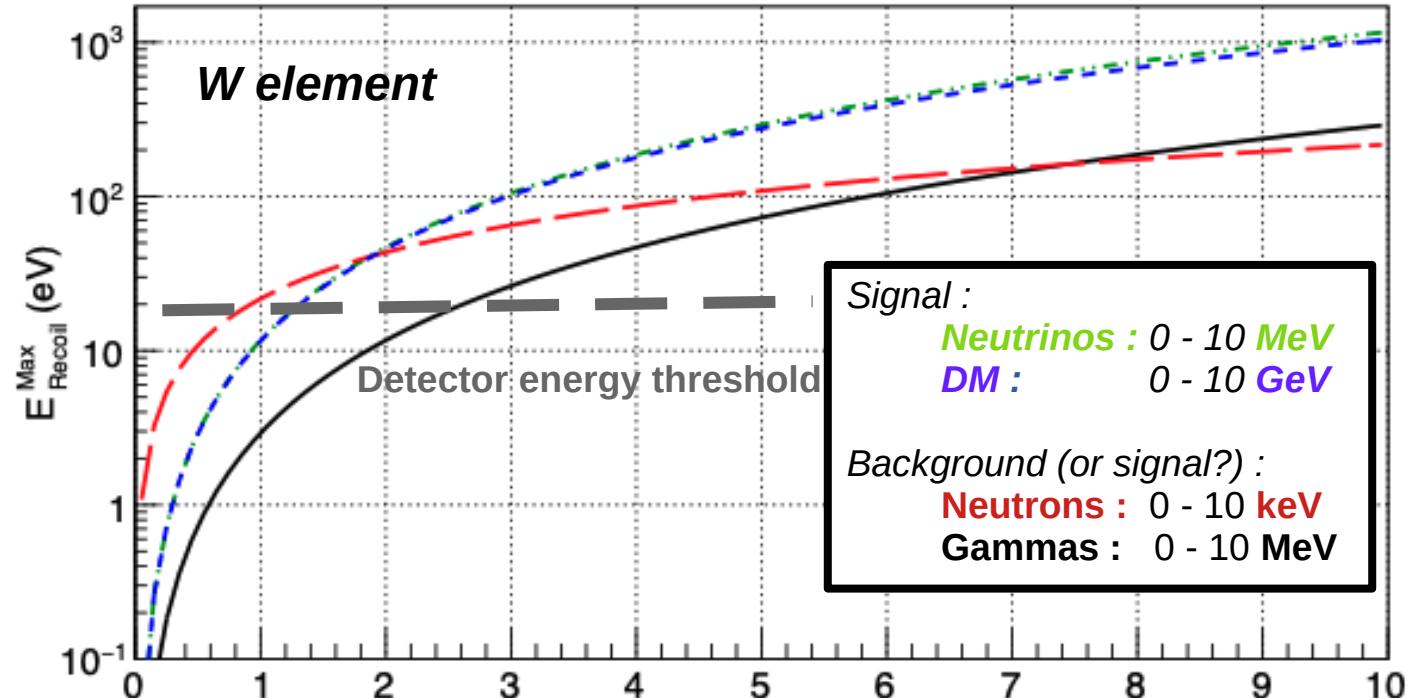
→ Impact of energy stored in lattice defect when reaching 100 eV scale ?
 ⇒ linearity study



NEED SUB-keV CALIBRATION METHODS

State-of-the-art calibration techniques :

- mainly electron recoils for *in-situ* calibration with LED [1], XRF source BUT surface calibration
- alphas BUT surface calibration
- epithermal/fast neutrons produced at accelerator AND limited by TOF and angular precisions



What about thermal neutrons ?

- gammas from (n,γ) reaction

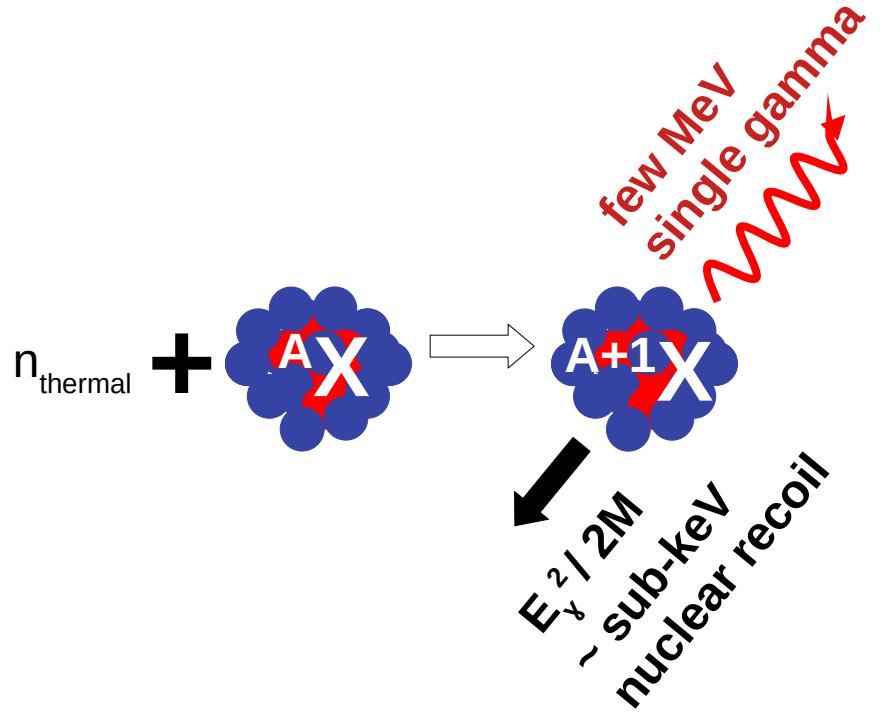
First indirect measurement by Jones and Kramer [2]

[1] L. Cardini et al., Eur. Phys. J. C 81 (2021) 7, 636.

[2] K.W. Jones and H.W. Kraner, Phys. Rev. A, 11 4, 1975

CRAB METHOD [1]

Thermal (~25meV) neutron radiative capture



The high-energy gamma leaves the cm scale detector without energy deposition

Advantages

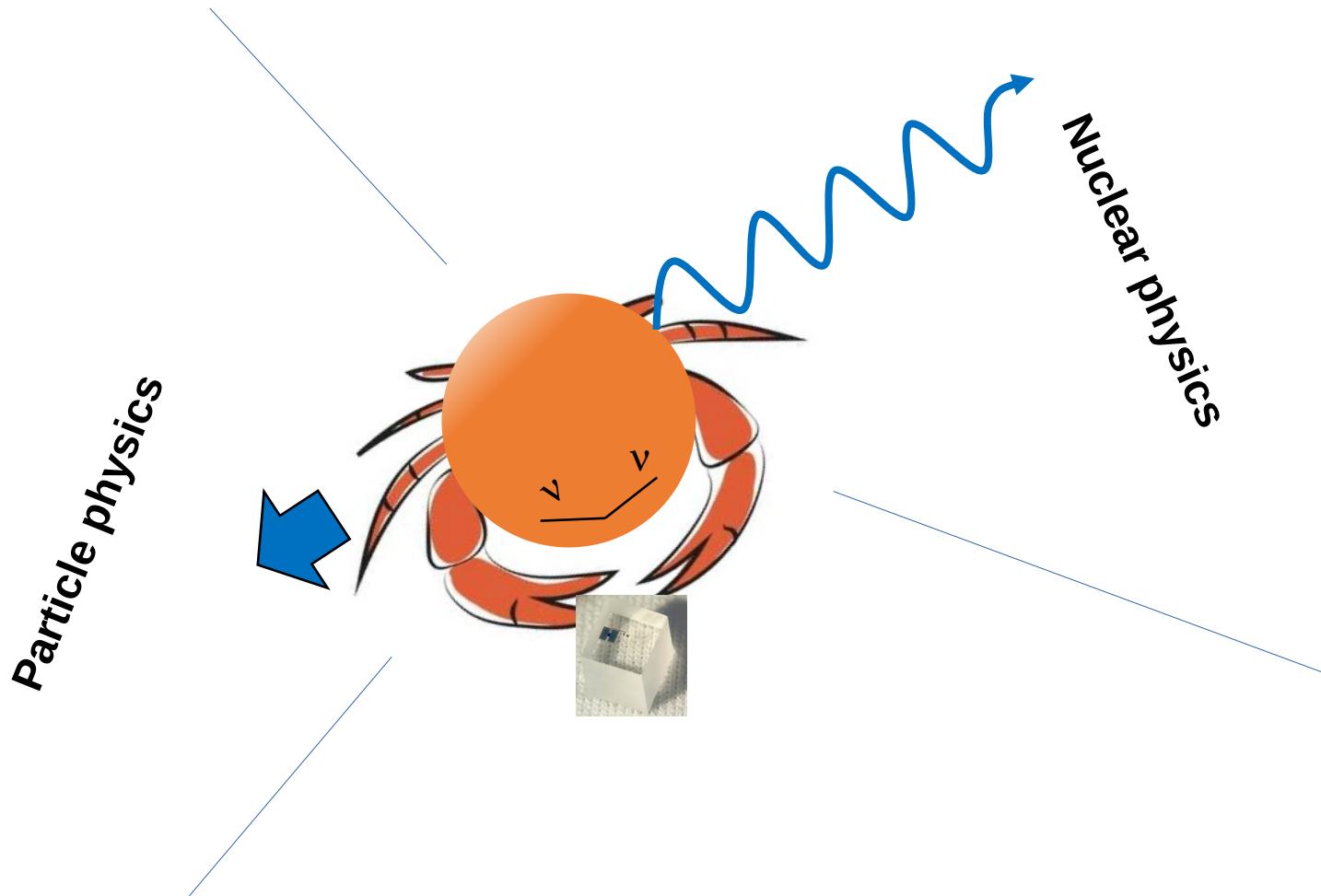
- Pure nuclear recoil \Rightarrow mimic the neutrino/DM signal
- Allows to probe the whole bolometer
- Accuracy \Rightarrow well defined peak

However non-trivial nucleus de-excitation to simulate

- transition probability from S_n to GS ?
 \Rightarrow signal intensity
- multi-gamma/electron cascade ?
 \Rightarrow background evaluation in the ROI
 \Rightarrow dead-time (typical response time of $\sim \text{ms}$ for cryogenic detectors)

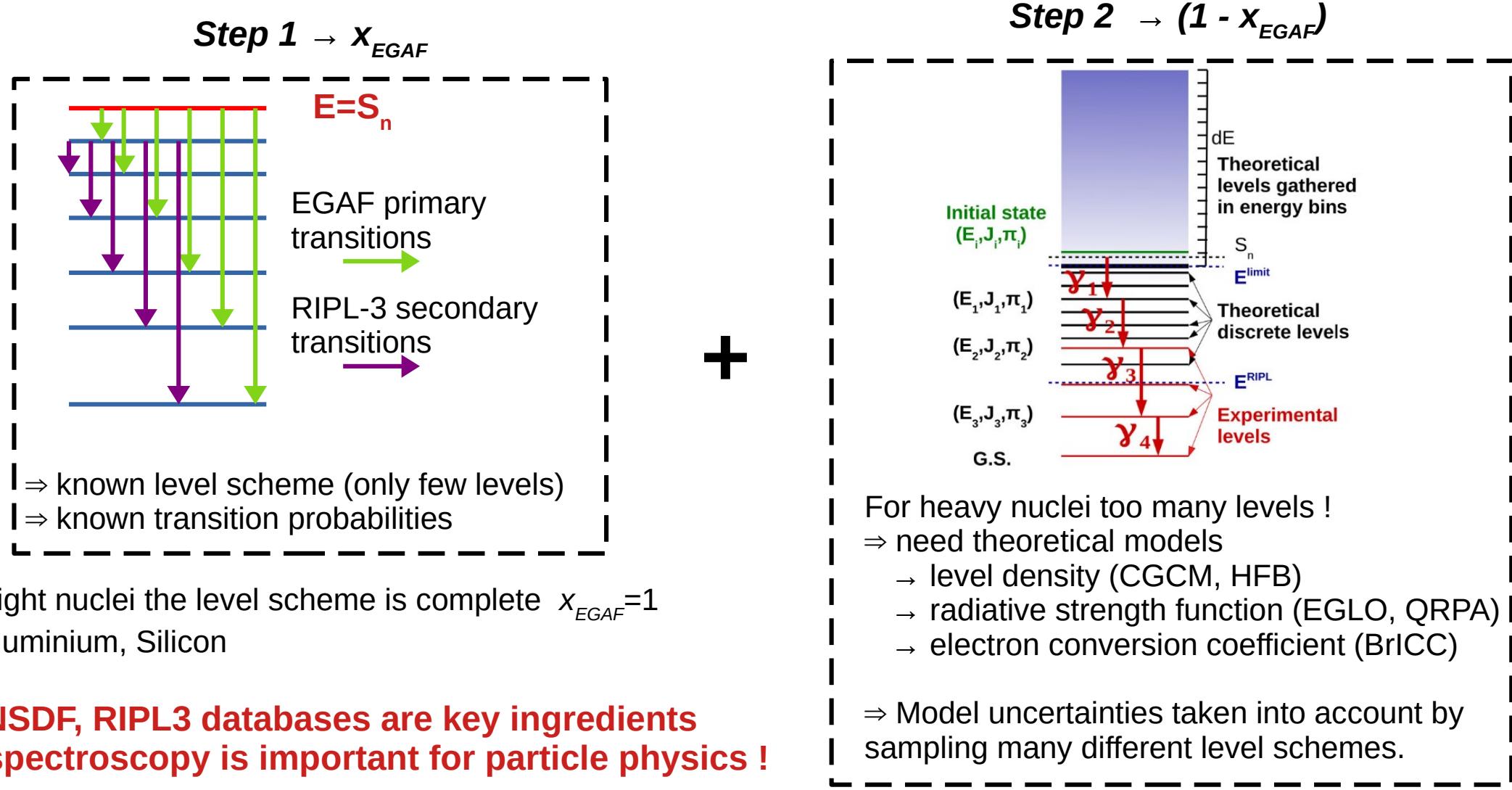


THE CRAB METHOD WHERE PARTICLE AND NUCLEAR PHYSICS MEET



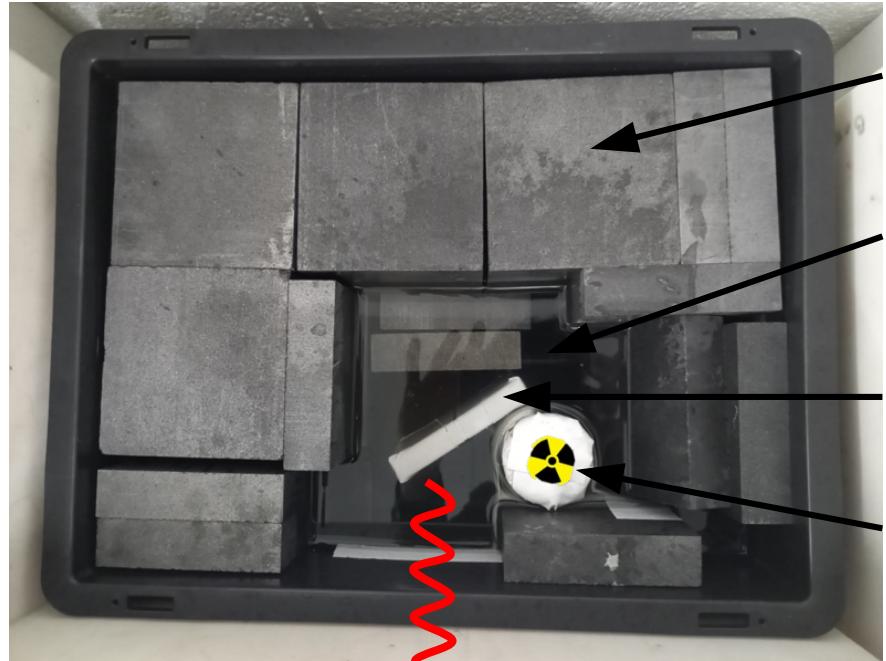
FIFRELIN SIMULATION

Fission fragment de-excitation code developed at CEA-Cadarache [1]

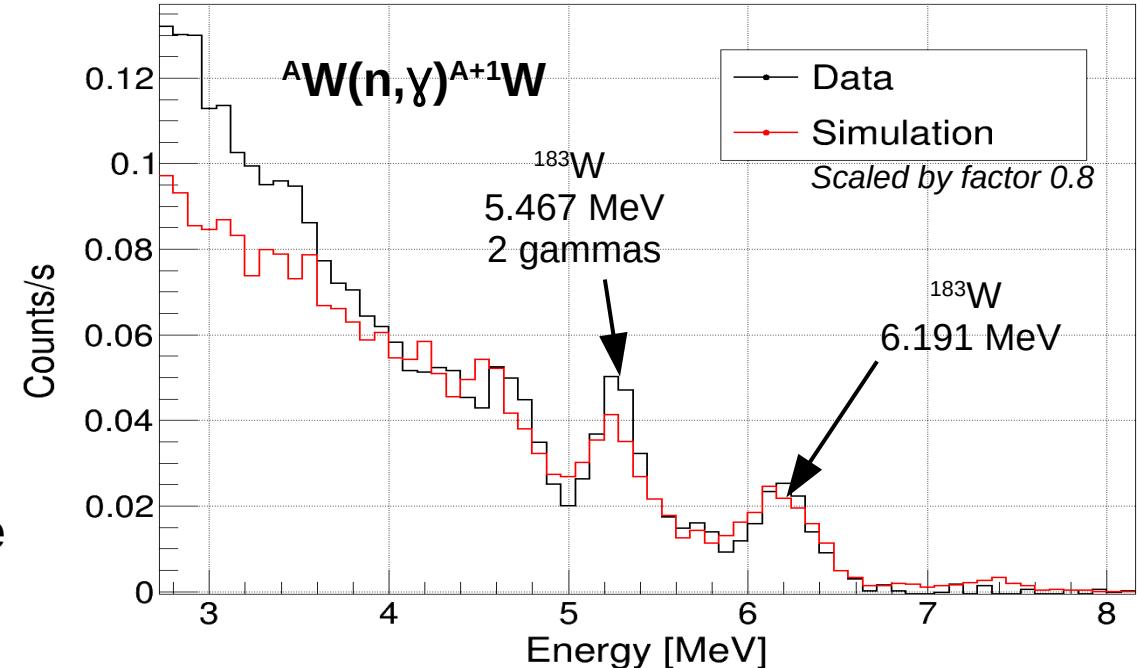


FIFRELIN VALIDATIONS – HOME MADE HIGH ENERGY GAMMA SOURCE

Already validated in the STEREO experiment on Gd isotopes [1,2]



Graphite reflector
Water moderator
Target
 (W, Ni, Fe, etc)
Neutron source
 ^{252}Cf



Discrepancies between data and MC less than 20 % for W isotopes
Confirmation of the FIFRELIN predictions : position and intensities of the S_n gamma line from tungsten.



STUDIED CRYO-DETECTOR MATERIALS

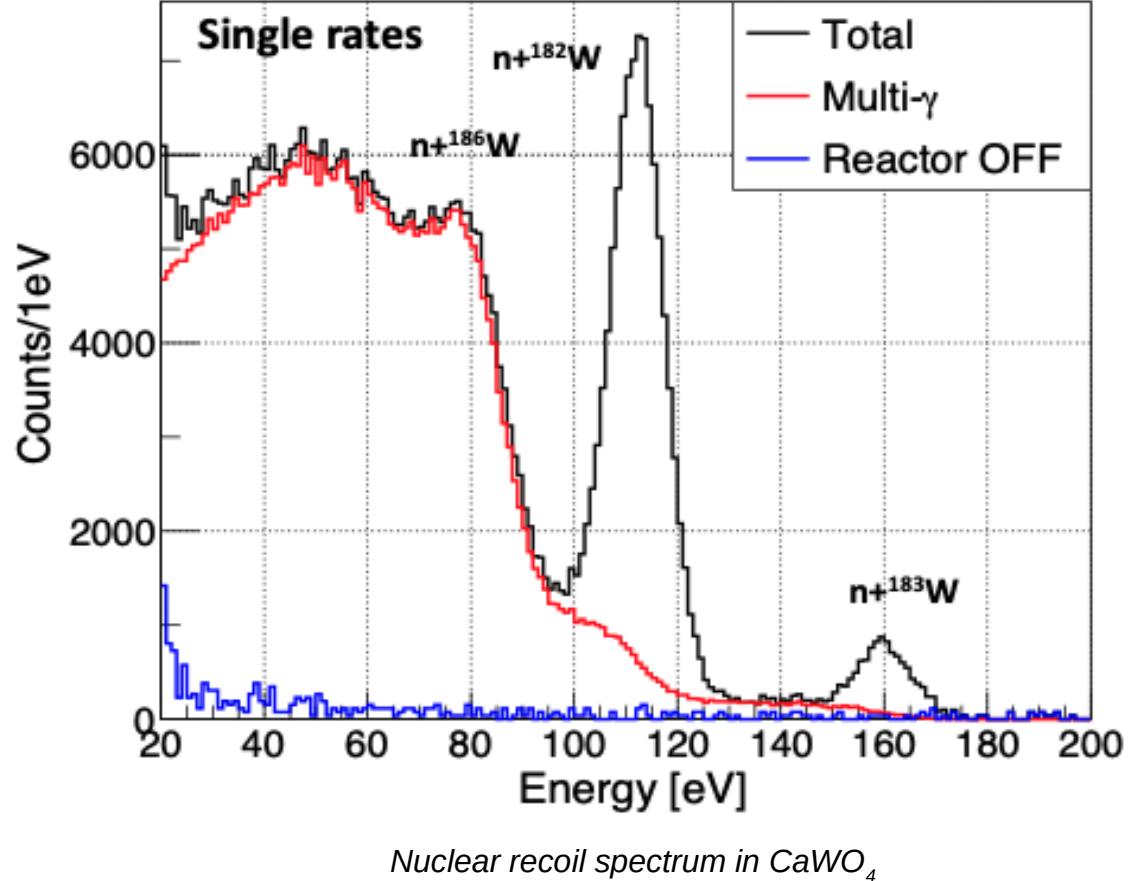
Crystal	Isotopes	Nucleus ${}^A X$		$+ n_{\text{thermal}}$	\rightarrow	Nucleus ${}^{A+1} X$		$+ g$
		Nat. ab. [%]	$\sigma_{(n,g)}$ [barn]			FOM	S_n [keV]	
CaWO₄ <i>CRESST</i> <i>NUCLEUS</i>	¹⁸² W	26.5	20.3	13.9	7478		6191	112.5
	¹⁸³ W	14.3	9.9	5.8	821		7411	160.3
	¹⁸⁴ W	30.6	1.6	1.5	73		5754	96.1
	¹⁸⁶ W	28.4	37.9	0.3	323		5467	85.8
Ge <i>EDELWEISS</i> <i>RICOCHET</i>	⁷⁰ Ge	20.5	3.1	2.0	127		7416	416.2
	⁷⁴ Ge	36.5	0.5	2.8	51		6506	303.2
Al₂O₃ <i>MINER</i> <i>NUCLEUS</i>	²⁷ Al	100	0.2	26.8	536		7725	1145
Si <i>SuperCDMS</i> <i>DAMIC SENSEI</i> <i>Skipper-CCD</i> <i>CONNIE</i>	²⁸ Si	92.2	0.2	2.2	41		8473	1330
		92.2	0.2	7.0	129		7199+1274	990
	²⁹ Si	4.7	0.1	6.7	3		10609	2016
	³⁰ Si	3.1	0.1	1.5	0.5		6587	752

Keep in mind :

- signal depends on detector resolution (σ)
⇒ high resolution needed

PREDICTED NUCLEAR RECOIL SPECTRA – CaWO₄

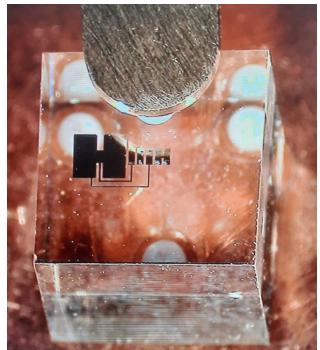
- Geant4 simulation based on TOUCANS [1] + FIFRELIN
- Mono-directionnal thermal neutron beam



- One prominent peak at 112 eV
⇒ could simply pop-up
- Second small peak at 160 eV
⇒ more difficult to get
- Third peak at ~85 eV
⇒ buried in multi-γ background
⇒ how to recover it ?
⇒ wait the end of the talk!



FIRST MEASUREMENT CRAB / NUCLEUS COLLABORATIONS



NUCLEUS CaWO₄
cryo-detector
 $E_{th} = 50$ eV
 $\sigma(E) = 6$ eV



Detector in a copper box spring
decoupled from cryostat vibration

cea



More copper to thermalize the
detector below 100 mK
TES transition ~10 mK



Thermal neutrons produced with a 3.7 MBq
 ^{252}Cf in a polyethylene and graphite moderator
⇒ $0.25 n_{th}/\text{s}$ on the cryo-detector

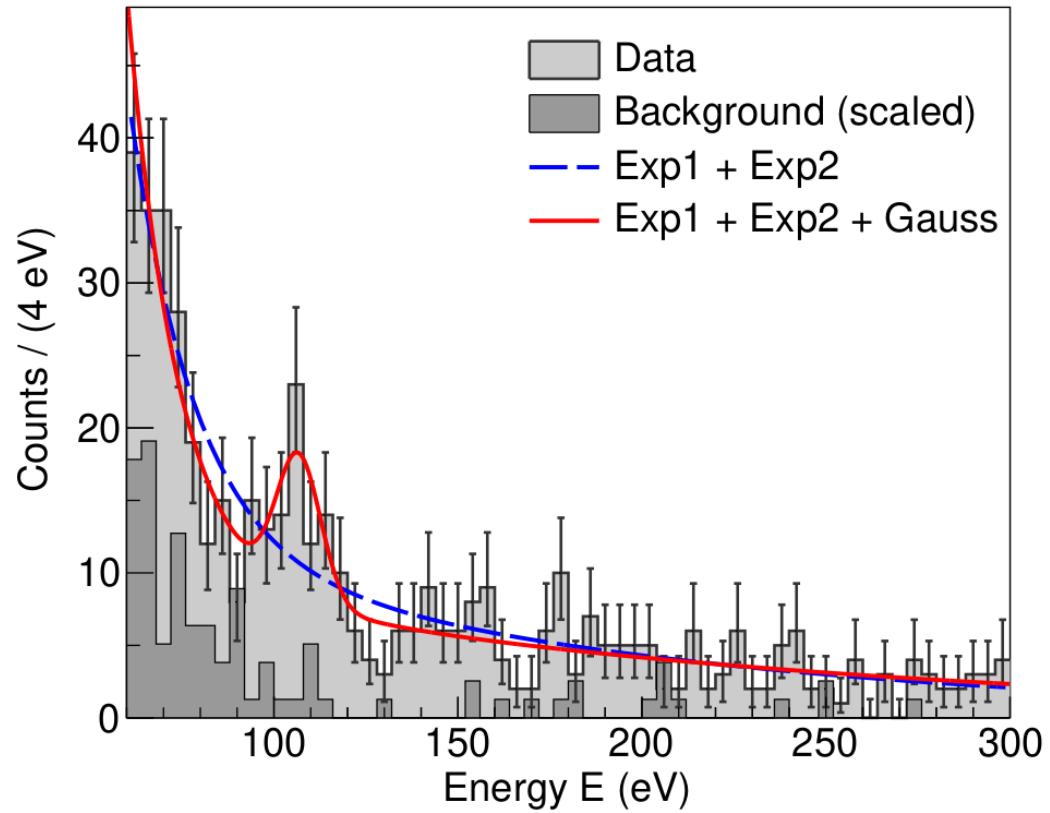
FIRST MEASUREMENT – RESULTS

Blind search peak

Test the presence of a peak

Background = 2 exponentials

Signal = gaussian



Presence of a peak
 $\Rightarrow 3.1\sigma$ significance (2-sided)

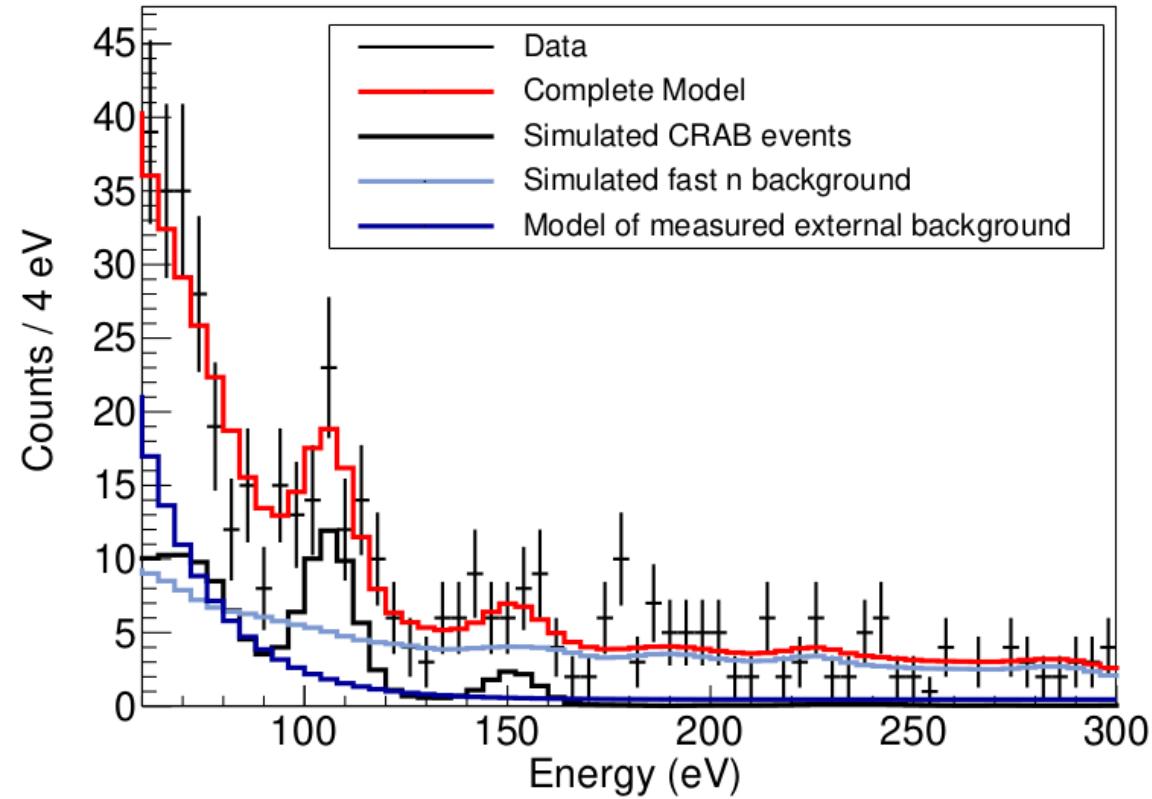
→ Background data lifetime = 18.9 h

→ Source data lifetime = 40.2 h

Test the presence of nuclear recoils from neutron capture

Background = exponential fit to bkgd data

Signal = GEANT4 + FIFRELIN



Presence of nuclear recoils from neutron capture
 $\Rightarrow 6\sigma$ significance (2-sided)
 $\Rightarrow \chi^2/NDF = 58.09/59$



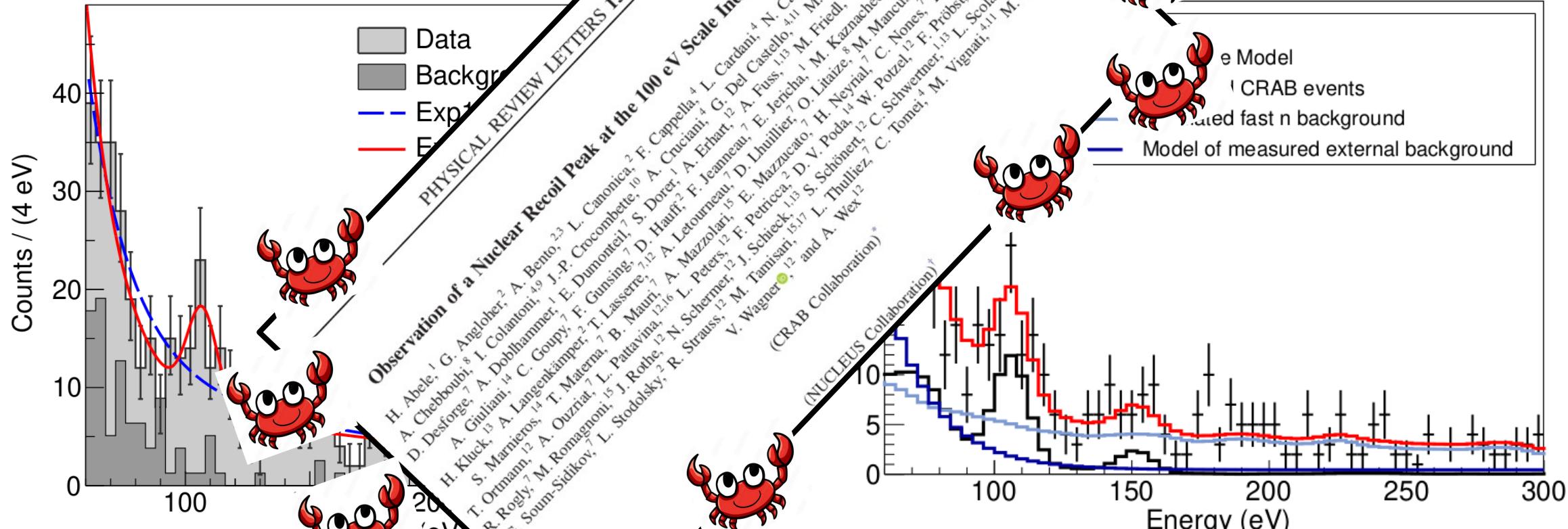
FIRST MEASUREMENT – RESULTS

Blind search peak

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Background = 2 exponentials

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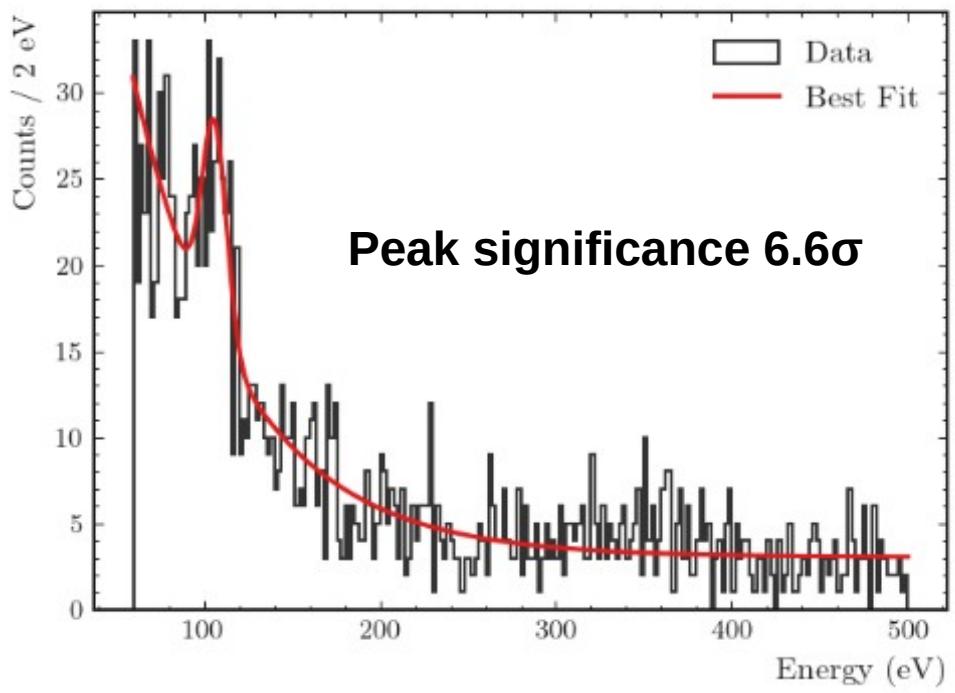




FIRST MEASUREMENT CONFIRMED BY OTHERS !

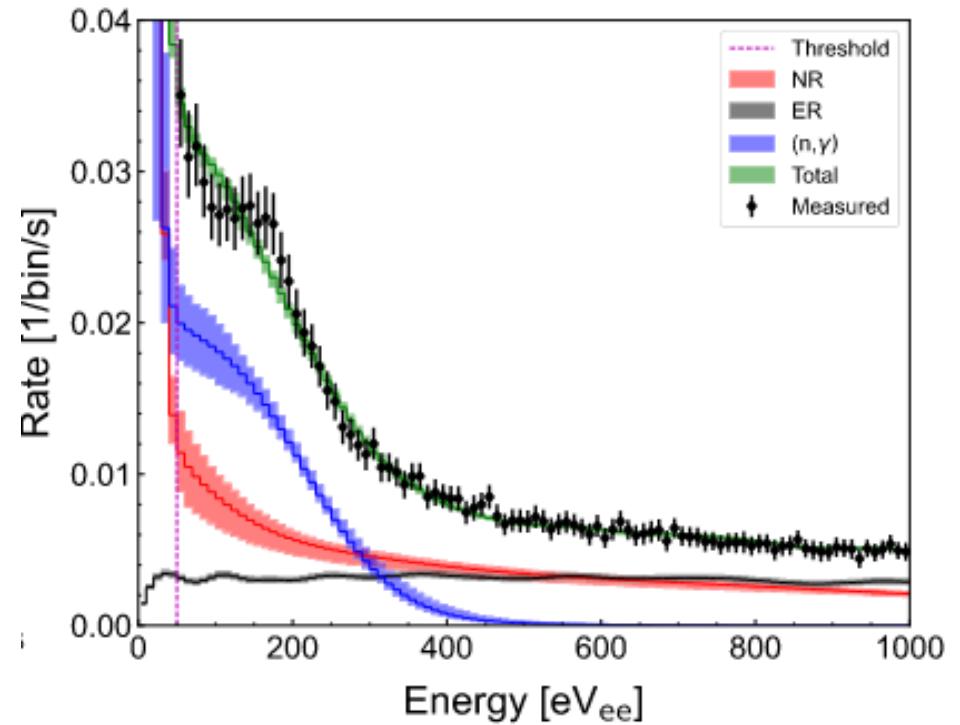
CRESST = dark matter with CaWO_4 cryo-detector
→ confirmation of our first CRAB signal with 3 detectors

<https://arxiv.org/pdf/2303.15315.pdf>
Accepted in Phys. Rev. D in June 2023



Super-CDMS = dark matter with Si cryo-detector
→ presence of nuclear recoils following thermal neutron captures

A. N. Villano, Phys. Rev. D, 105, 083014, 2022

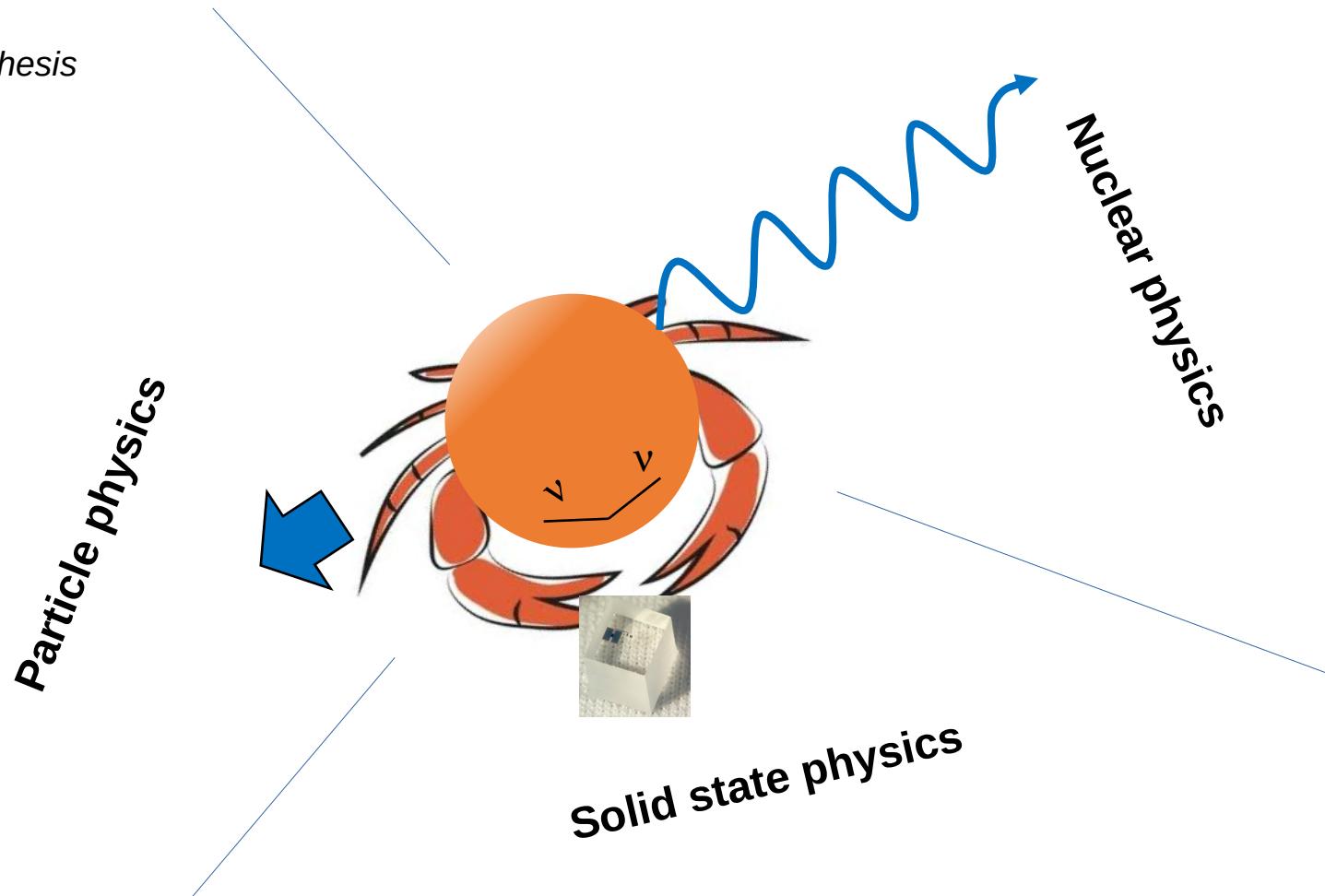


CRAB has already a big impact in the dark matter and neutrino communities !



TIMING EFFECT WHERE NUCLEAR AND SOLID STATE PHYSICS MEET

Gabrielle Soum-Sidikov thesis

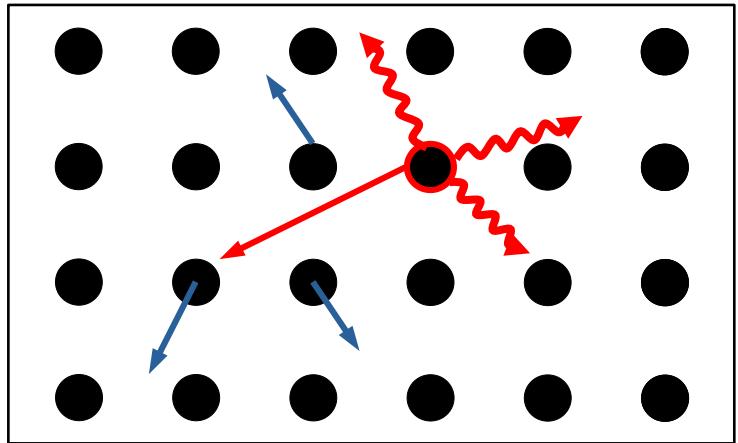




TIMING EFFECT – EXTREME HYPOTHESES

Fast hypothesis

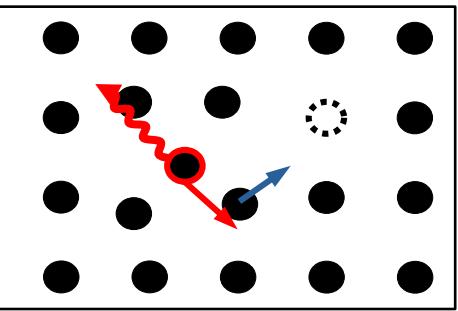
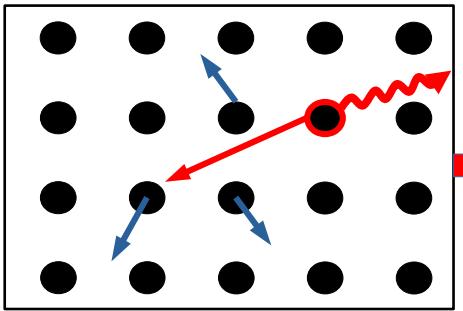
$$\tau_{recoil} \gg \tau_{cascade} \quad E_{recoil} = (\sum_i \vec{p}_i)^2 / 2M$$



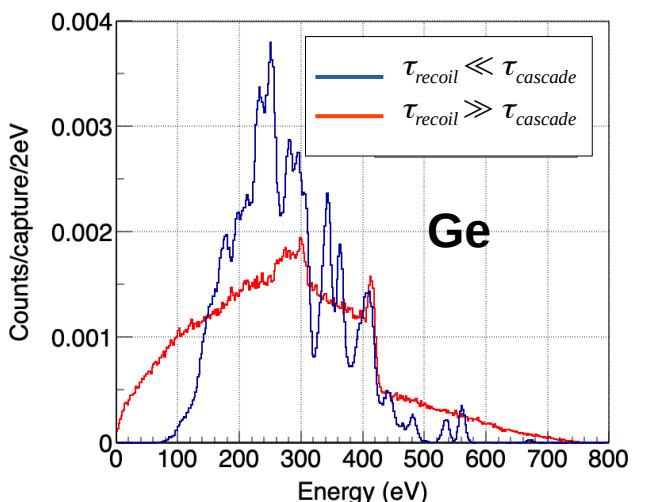
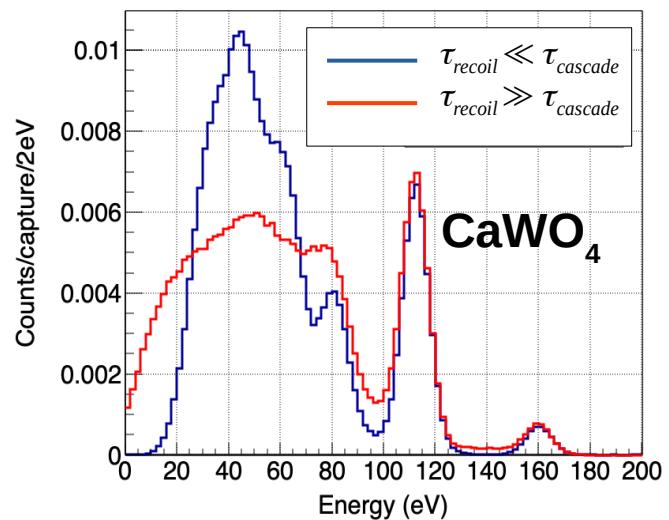
Reality in between ?

Slow hypothesis

$$\tau_{recoil} \ll \tau_{cascade} \quad E_{recoil} = \sum_i E_{recoil,i} = \sum_i p_i^2 / 2M$$



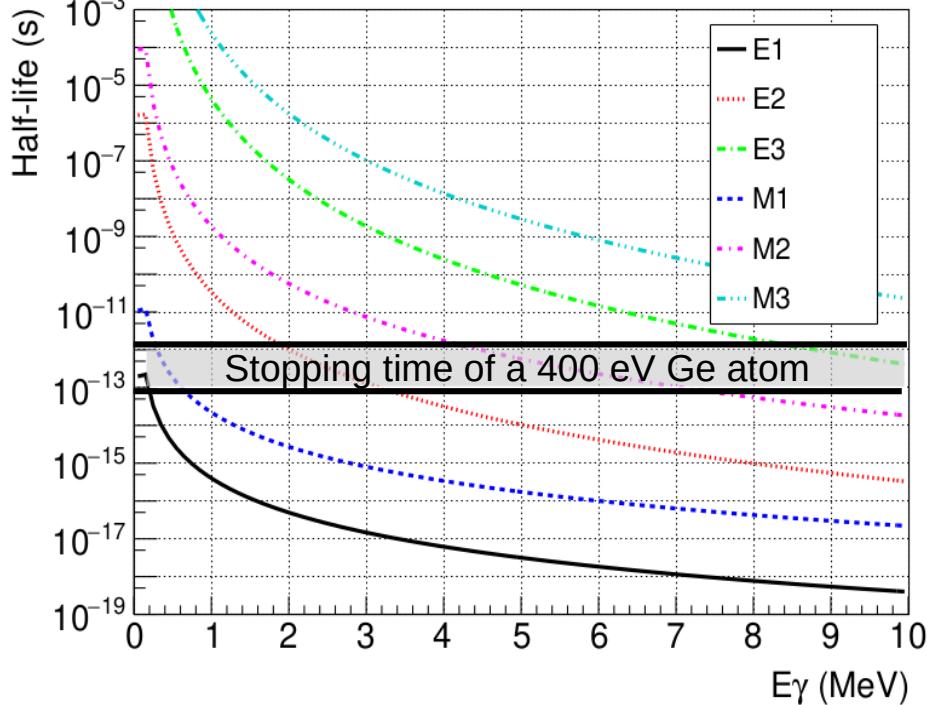
etc ...





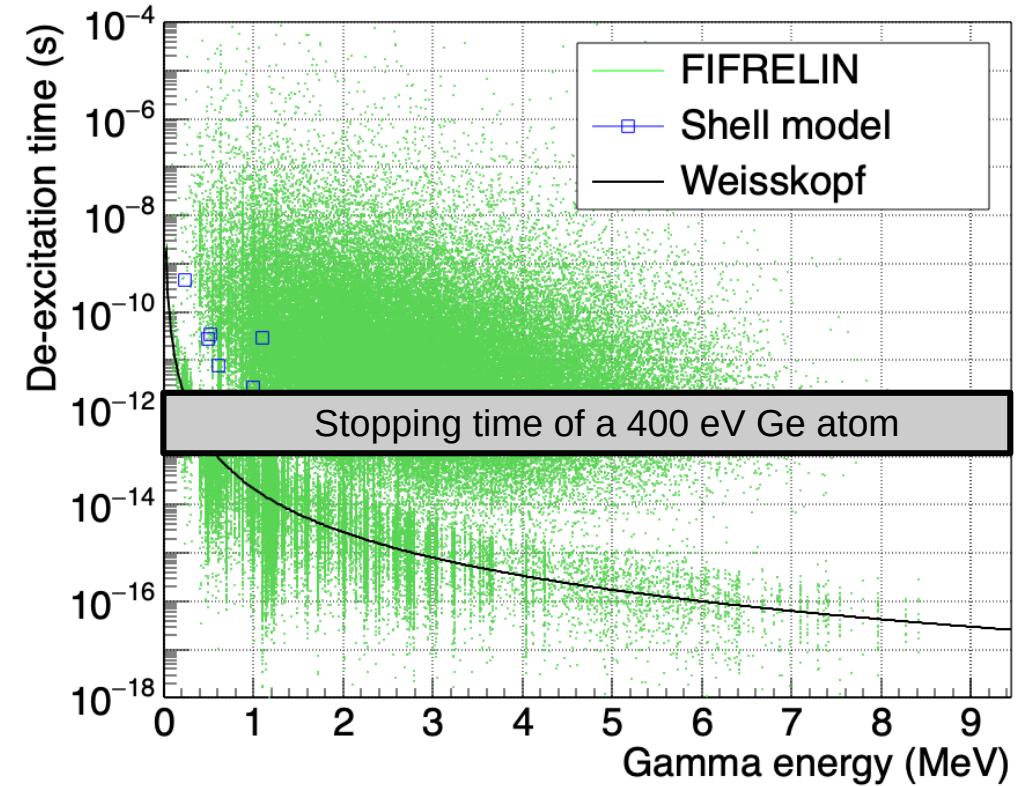
GAMMA TIMING FROM RADIATIVE PARTIAL WIDTH Γ_γ

Γ_γ from Weisskopf estimate
→ single particle approximation



→ known to be inaccurate by a factor $\sim 10 - 100$
⇒ no collective effects

Γ_γ with collective effect



FIFRELIN Γ_γ computed with

- EGLO radiative strength function
- CGCM nuclear level density

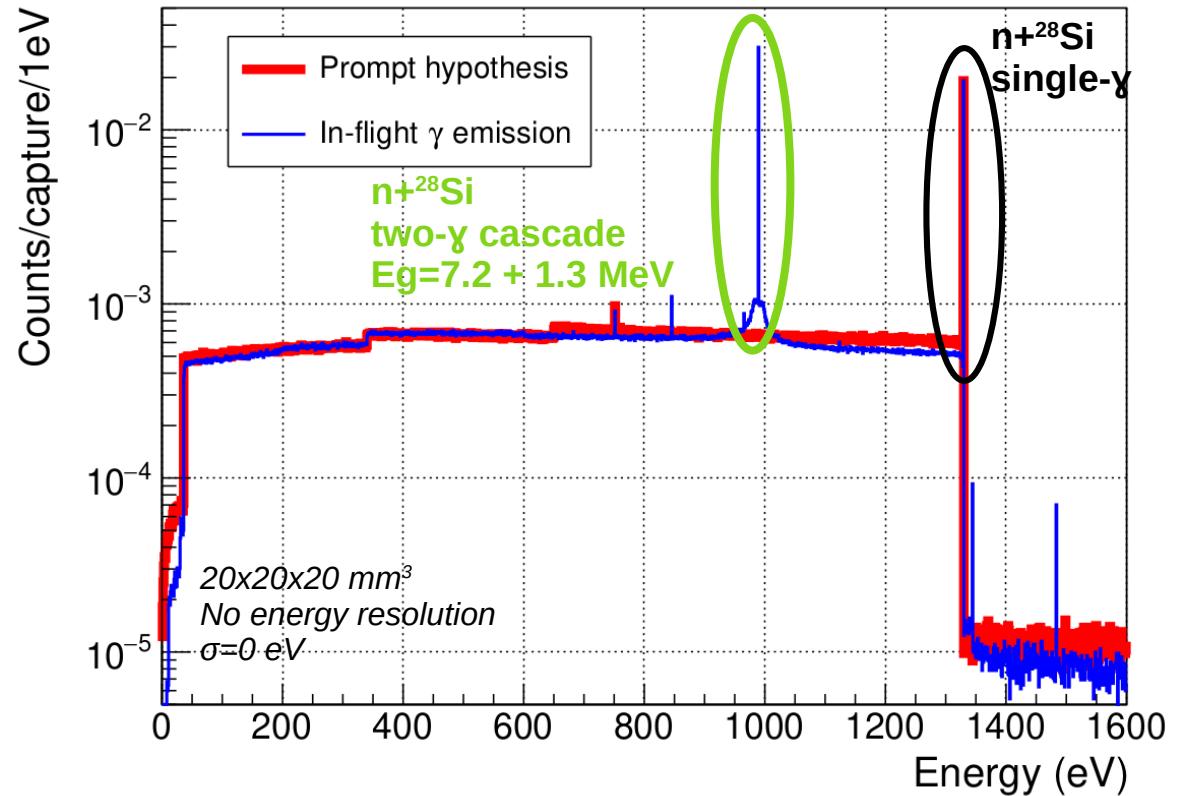
Effective way
(parametrisation) to take
collective effects into
account

Shell model : Kshell code + JUN45 interaction

G. Soum-Sidikov,¹ H. Abele,² J. Burkhart,³ F. Cappella,⁴ N. Casali,⁴ R. Cerulli,^{5,6} A. Chalil,^{1,7} A. Chebboubi,⁷ J-P. Crocombette,⁸ G. del Castello,^{9,10} M. del Gallo Roccagiovine,^{9,10} A. Doblhammer,² S. Dorer,² E. Dumonteil,¹ A. Erhart,¹¹ A. Giuliani,¹² C. Goupy,¹ F. Gunsing,¹ E. Jericha,² M. Kaznacheeva,¹¹ A. Kinast,¹¹ H. Kluck,³ A. Langenkämper,¹¹ T. Lasserre,^{1,11} A. Letourneau,¹ D. Lhuillier,¹ O. Litaize,⁷ P. de Marcillac,¹² S. Marnieros,¹² R. Martin,¹ T. Materna,¹ E. Mazzucato,¹ C. Nones,¹ T. Ortmann,¹¹ L. Pattavina,^{6,13} D.V. Poda,¹² L. Peters,¹¹ J. Rothe,¹¹ N. Schermer,¹¹ J. Schieck,^{2,3} S. Schönert,¹¹ O. Serot,⁷ L. Stodolsky,¹⁴ R. Strauss,¹¹ L. Thulliez,¹ M. Vignati,^{9,10} M. Vivier,¹ V. Wagner,¹¹ and A. Wex¹¹

(CRAB Collaboration)

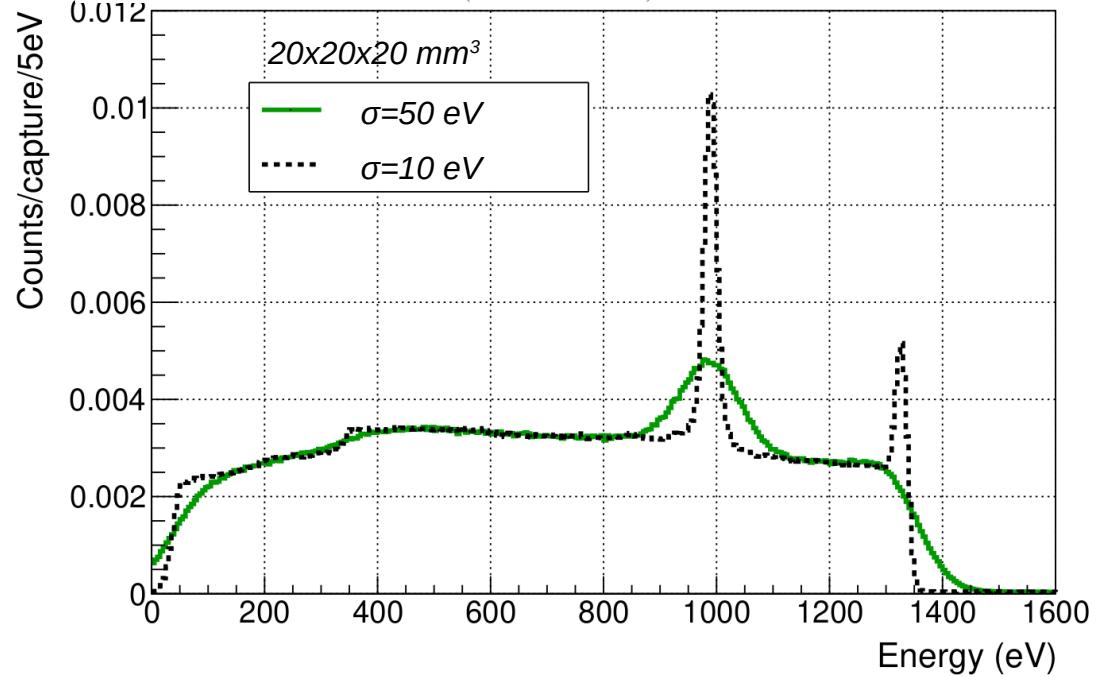
TIMING EFFECTS – SILICON



Most of the half-lives are in the databases
(collective effects taken into account)

In-flight gamma emission

- ⇒ probe nucleus de-excitation time
- ⇒ probe inter-atomic potentials
- ⇒ do you remember with GRID at GAMS @ILL [1] ?

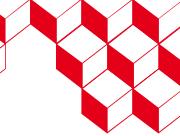


Two- γ cascade robust against poorer energy resolution

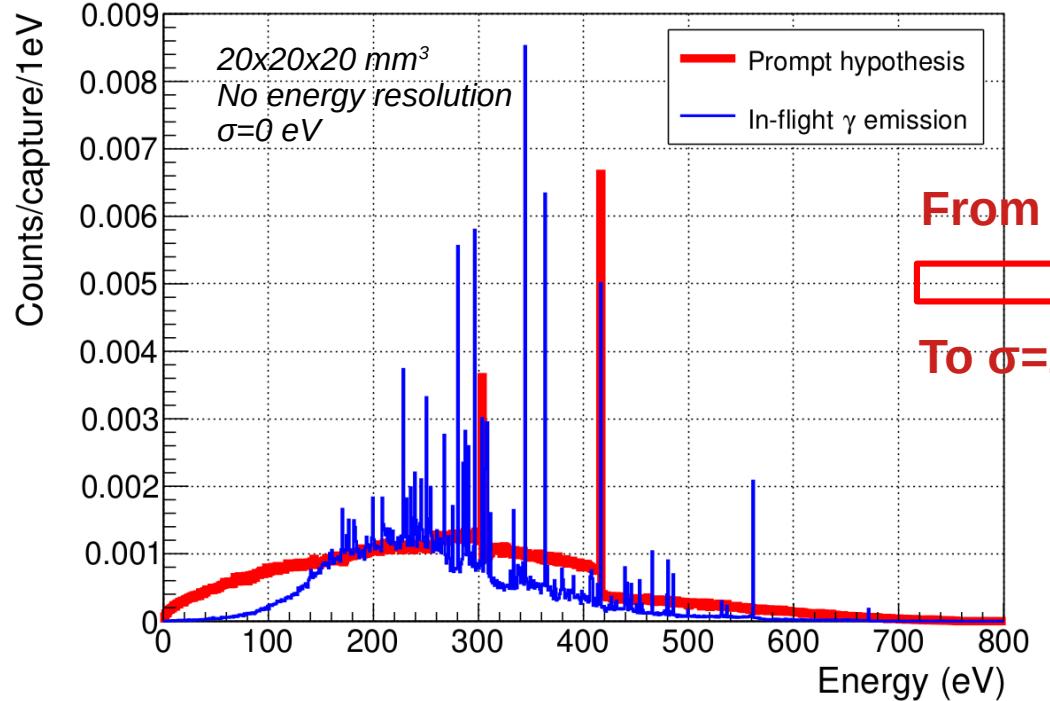
[0] G. Soum-Sidikov, Study of collision and γ -cascade times following neutron-capture processes in cryogenic detectors, <https://arxiv.org/abs/2305.10139> Submitted to Phys. Rev. D

[1] E. G. Kessler, Nucl. Instrum. Meth. A 457 (2001) 187:202

TIMING EFFECTS – GERMANIUM

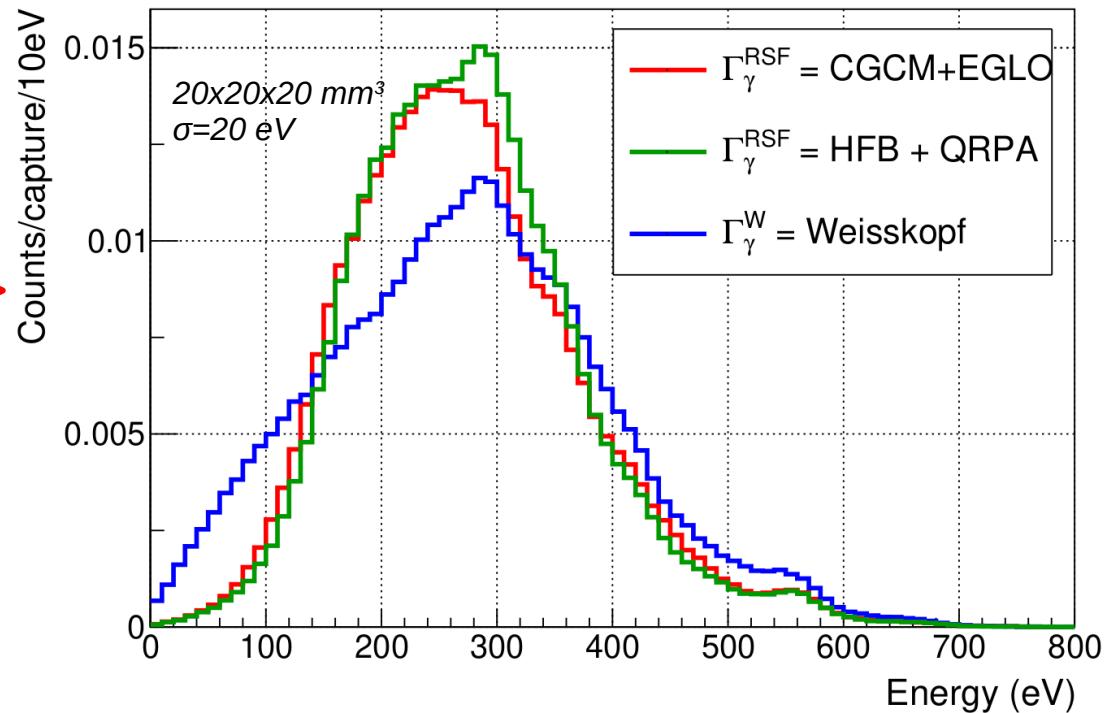


G. Soum-Sidikov,¹ H. Abele,² J. Burkhart,³ F. Cappella,⁴ N. Casali,⁴ R. Cerulli,^{5,6} A. Chalil,^{1,7} A. Chebboubi,⁷ J-P. Crocombette,⁸ G. del Castello,^{9,10} M. del Gallo Roccagiovine,^{9,10} A. Doblhammer,² S. Dorer,² E. Dumonteil,¹ A. Erhart,¹¹ A. Giuliani,¹² C. Goupy,¹ F. Gunsing,¹ E. Jericha,² M. Kaznacheeva,¹¹ A. Kinast,¹¹ H. Kluck,³ A. Langenkämper,¹¹ T. Lassere,^{1,11} A. Letourneau,¹ D. Lhuillier,¹ O. Litaize,⁷ P. de Marcillac,¹² S. Marnieros,¹² R. Martin,¹ T. Materna,¹ E. Mazzucato,¹ C. Nones,¹ T. Ortmann,¹¹ L. Pattavina,^{6,13} D.V. Poda,¹² L. Peters,¹¹ J. Rothe,¹¹ N. Schermer,¹¹ J. Schieck,^{2,3} S. Schönert,¹¹ O. Serot,⁷ L. Stodolsky,¹⁴ R. Strauss,¹¹ L. Thulliez,¹ M. Vignati,^{9,10} M. Vivier,¹ V. Wagner,¹¹ and A. Wex¹¹
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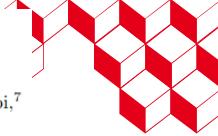
In-flight γ emission \Rightarrow more calibration peaks

Resolution is a critical parameter !



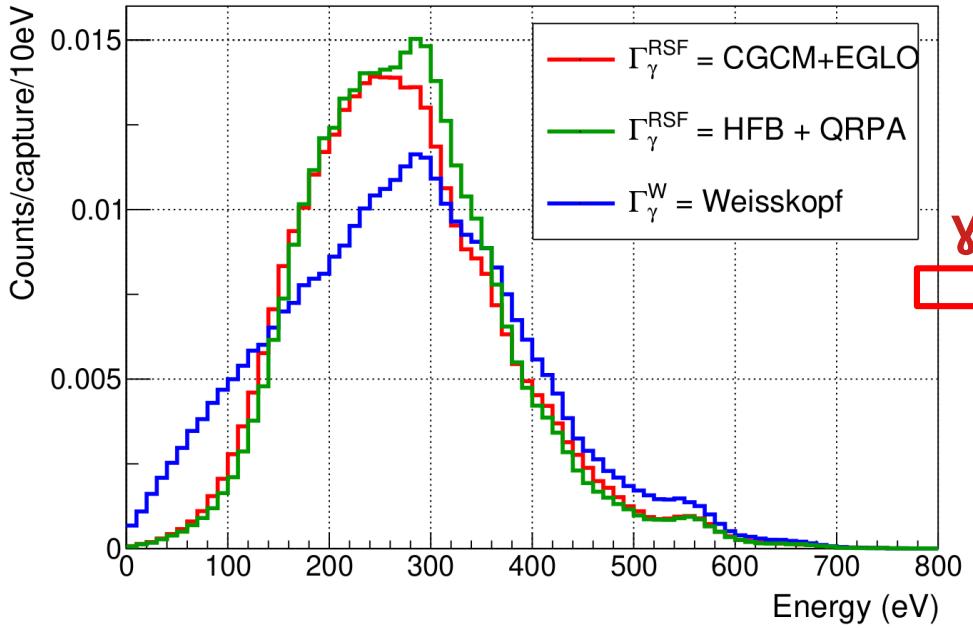
Recoil energy spectrum sensitive to nuclear models !
 \Rightarrow could help set constraints on models

[0] G. Soum-Sidikov, Study of collision and γ -cascade times following neutron-capture processes in cryogenic detectors, <https://arxiv.org/abs/2305.10139> Submitted to Phys. Rev. D

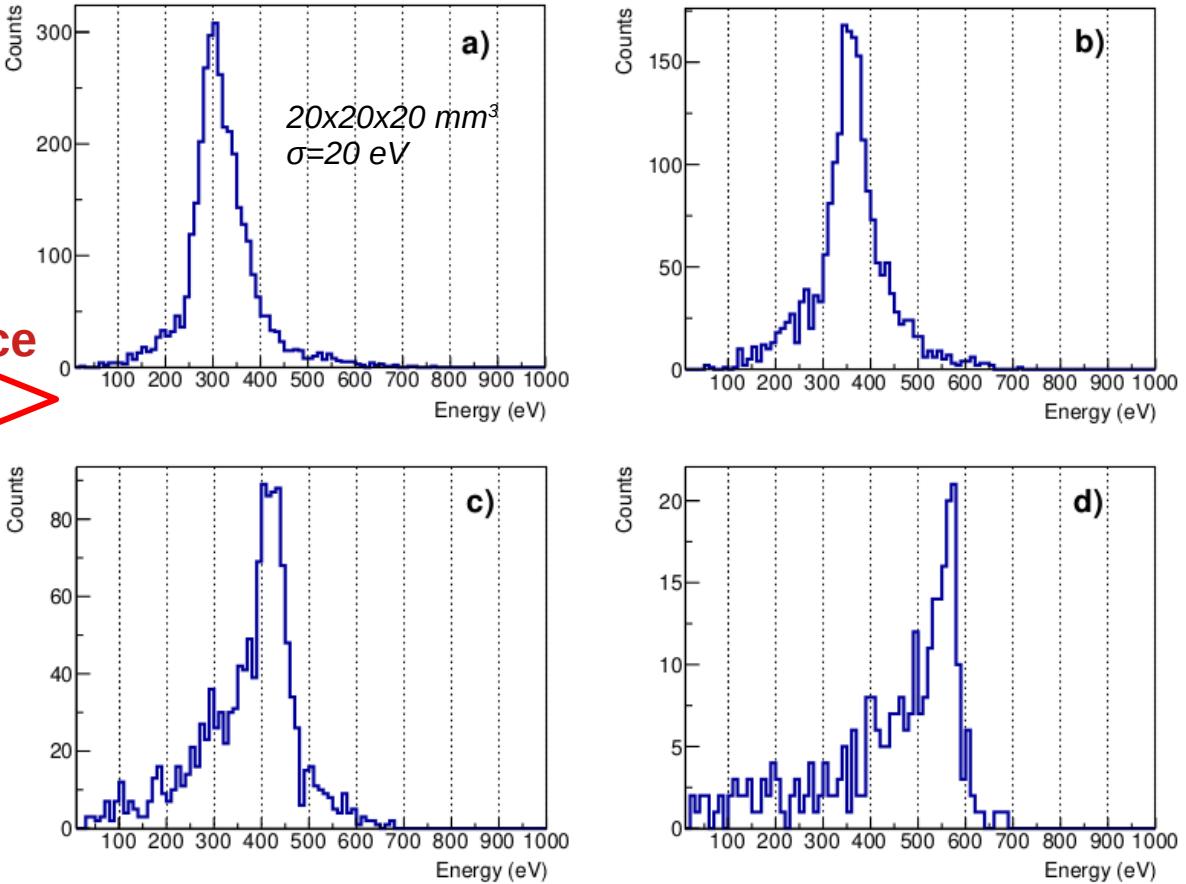


TIMING EFFECTS – GERMANIUM

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(CRAB Collaboration)



γ -coincidence



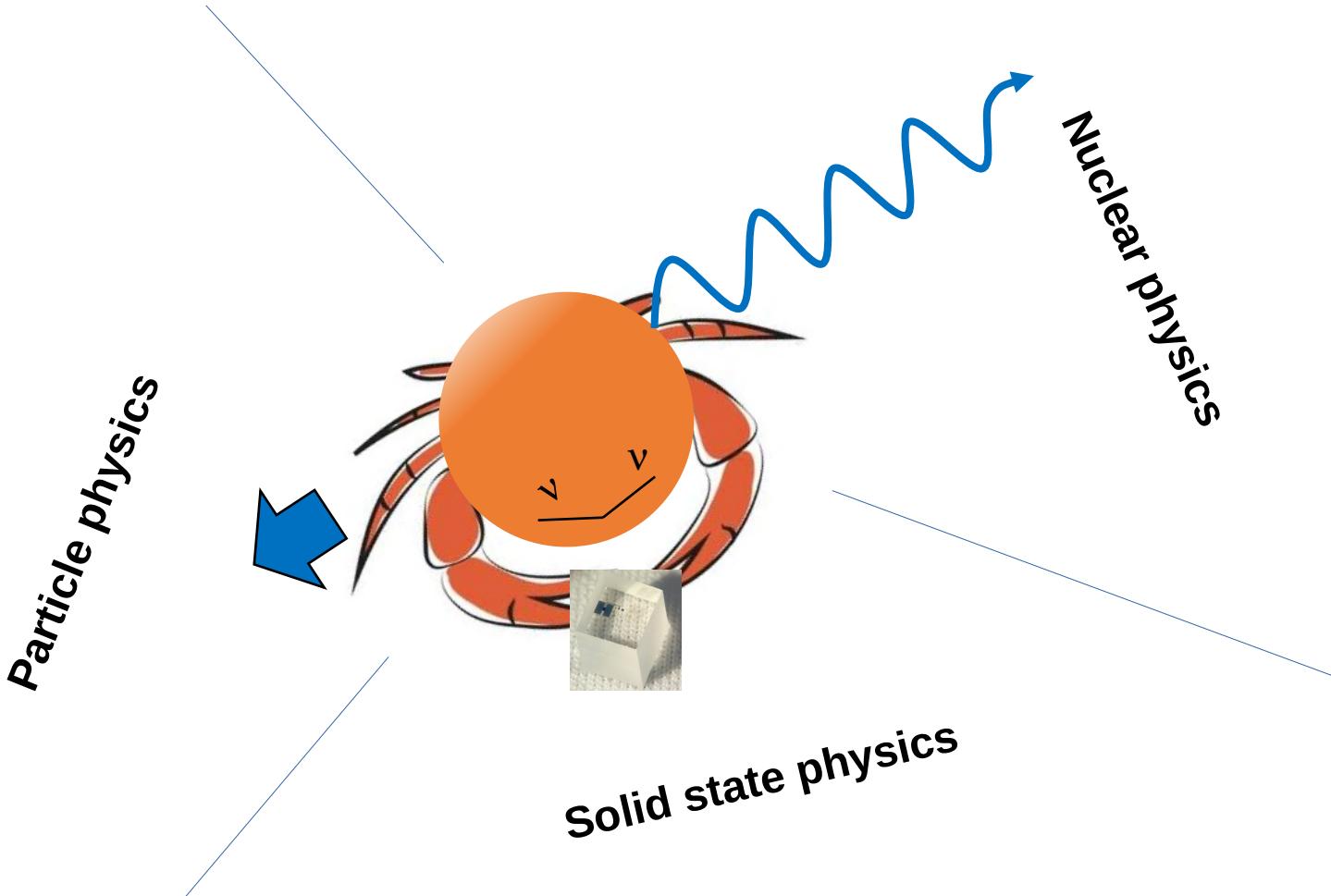
Resolution is a critical parameter

BUT gamma coincidence allows to overcome this limitation !

[0] G. Soum-Sidikov, Study of collision and γ -cascade times following neutron-capture processes in cryogenic detectors, <https://arxiv.org/abs/2305.10139> Submitted to Phys. Rev. D

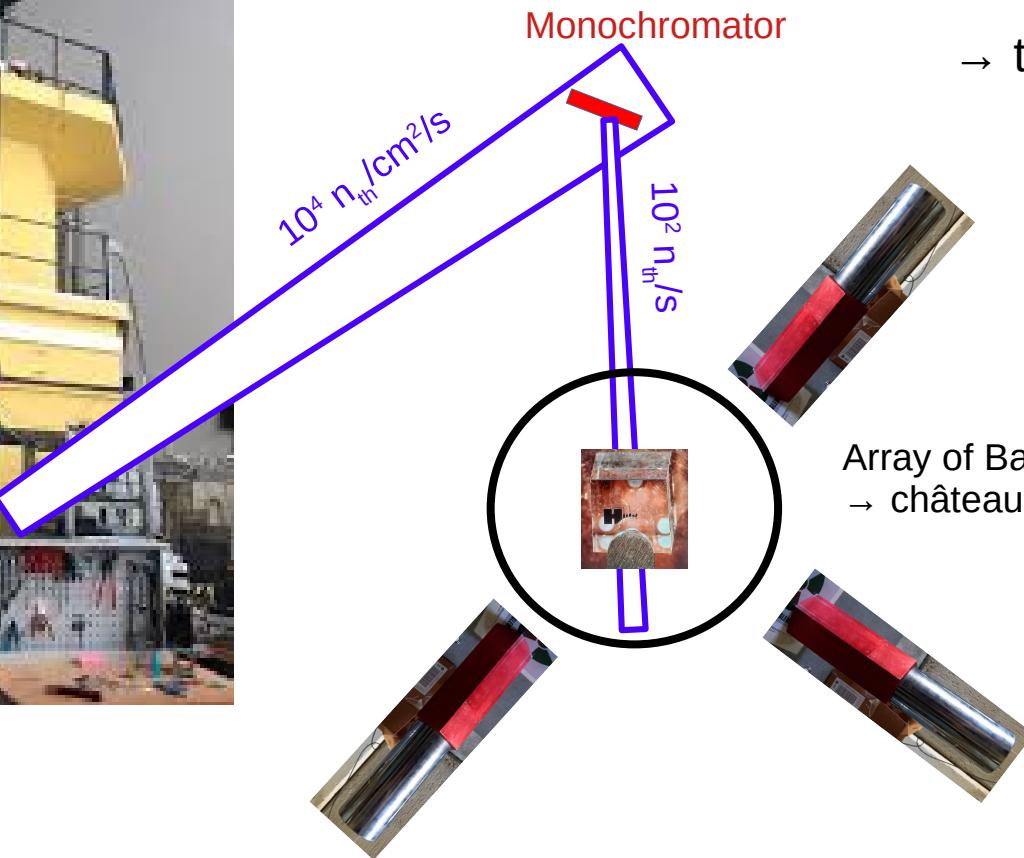


WHAT IS NEXT ?



HIGH PRECISION MEASUREMENTS – SOON IN 2024

Low intensity ($\sim 100 \text{ n}_{\text{th}}/\text{s}$) thermal neutron beam at TRIGA Mark-II nuclear reactor (250 kW) in Vienna



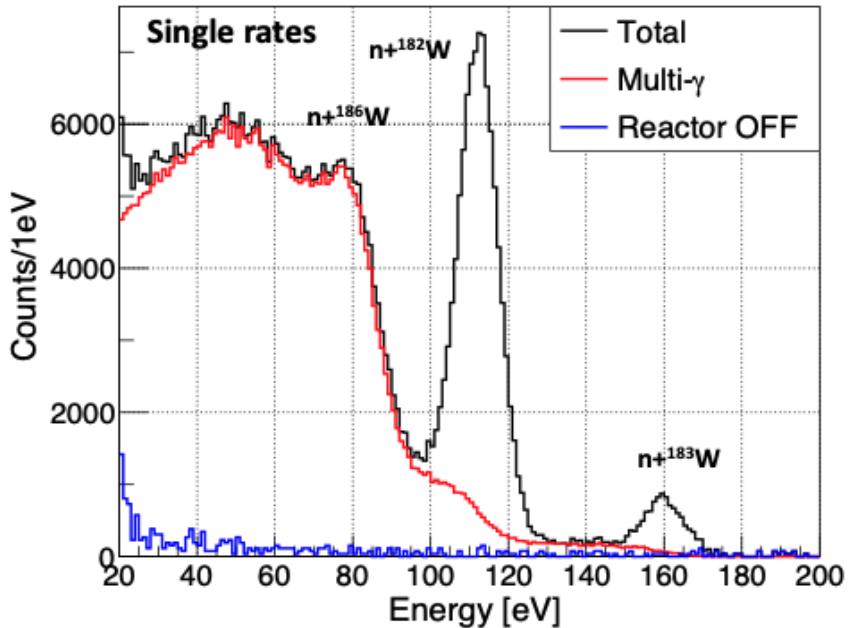
Coincidence between γ detectors and cryo-detector
→ increase signal / background ratio
→ get more calibration peaks
⇒ linearity study
→ test directionnality effect

Array of BaF₂ gamma detectors
→ château de cristal



DO YOU REMEMBER SLIDE 14 ?

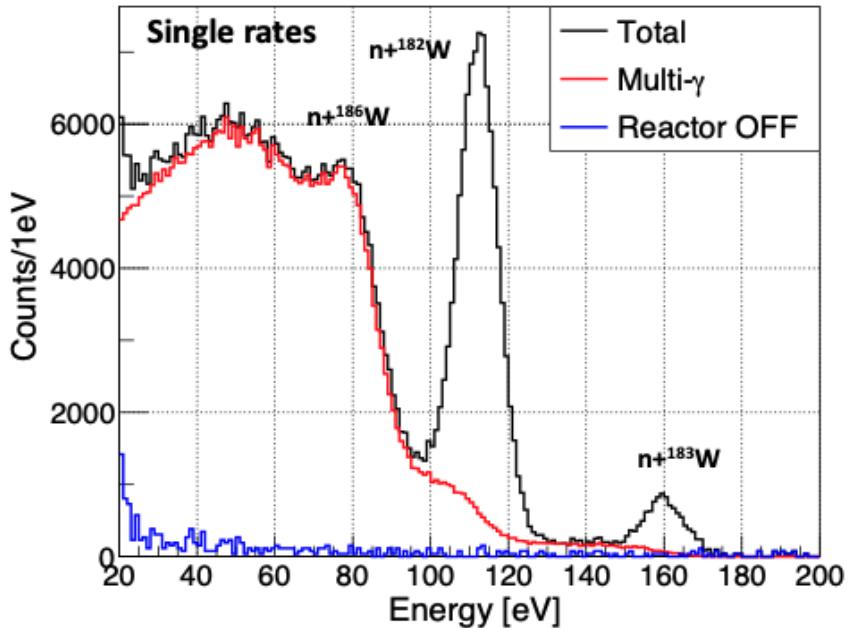
How to get a third peak at ~85 eV from the spectrum below ?



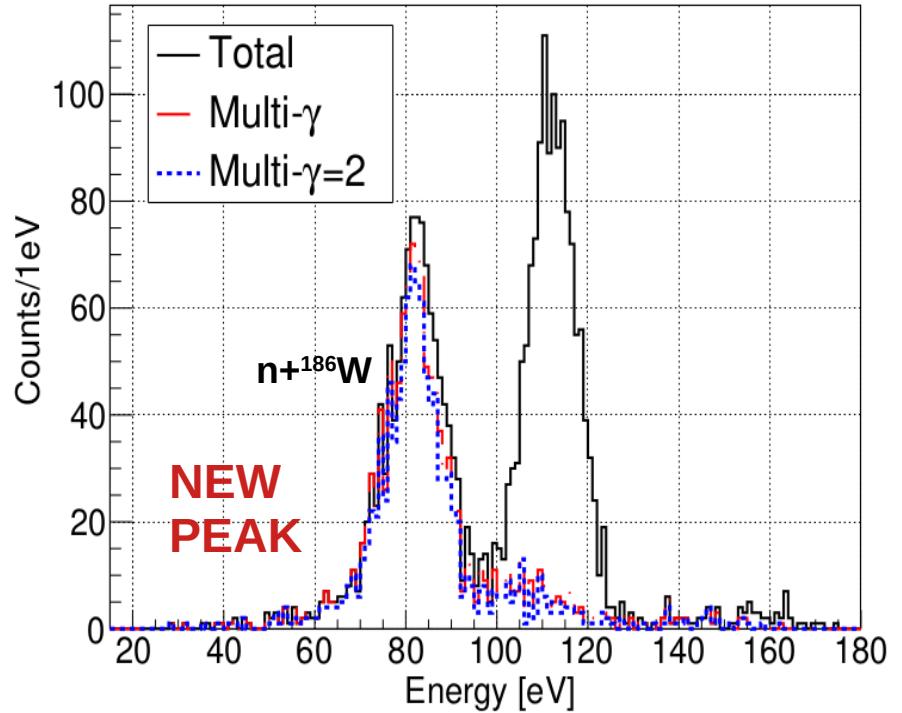


DO YOU REMEMBER SLIDE 14 ?

How to get a third peak at ~85 eV from the spectrum below ?



γ -coincidence !



*Recoil spectrum after gamma tagging with
 $E_\gamma = 5.5 \pm 0.2 \text{ MeV}$*

γ -tagging is a powerful tool to get a 3rd peak for CaWO₄ ⇒ **linearity study** !



CONCLUSION AND PERSPECTIVES

- **CRAB method** promising for a **sub-kev calibration** of the majority of cryo-detector materials in DM/CEvNS communities currently used (**CaWO₄, Ge, Si, Al₂O₃**)
- Successfull first measurement with a NUCLEUS CaWO₄ and a portable neutron source
 - ⇒ presence of a peak at ~112 eV with 3.1σ significance
 - Confirmation from independent measurement by the CRESST collaboration with higher significance
 - ⇒ presence of nuclear recoils with 6σ significance
- CRAB with **gamma tagging** is a powerful tool to **increase S/B** and **access lower energy** recoils, study the linearity of the bolometer response and tag the direction of the recoil (directionality)
- CRAB phase 2 / full precision in Vienna in preparation should be performed in 2024
 - ⇒ stay tuned !



irfu

THANK YOU



The CRAB collaboration